

# Particle Physics II

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## Exercise sheet VI

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### In-class exercises

#### **Exercise 36** *Recapitulation*

The following are only suggestions; we can discuss selected topics from below and/or any other topics.

- (i) Helicity, chirality: How are they defined, what are their properties, why are they important?
- (ii) V-A theory: What does it mean; and how can it be used to describe an (electro-weak) process?
- (iii) Goldhaber experiment: How does it work?
- (iv) Pion decay: Why are muons preferred?
- (v) ...

## Home exercises

### Exercise 37 Muon decay

10 Punkte

Consider the decay of a muon,

$$\mu^-(p_1) \rightarrow e^-(p_4) + \bar{\nu}_e(p_2) + \nu_\mu^-(p_3).$$

- (i) Draw the Feynman diagram for the leading-order contribution to this process.  
 (ii) Show that in the context of the V-A theory description of leptonic currents, the decay amplitude can be written as

$$\mathfrak{M} = \frac{G_F}{\sqrt{2}} [\bar{u}_3 \gamma^\mu (1 - \gamma^5) u_1] [\bar{u}_4 \gamma_\mu (1 - \gamma^5) v_2]$$

with  $G_F = \frac{\sqrt{2}}{8} \left( \frac{g_W}{M_W} \right)^2$ .

- (iii) Using the relation

$$\text{Tr} \left[ \gamma^\mu (1 - \gamma^5) \not{p}_1 \gamma^\nu (1 - \gamma^5) \not{p}_2 \right] \times \text{Tr} \left[ \gamma_\mu (1 - \gamma^5) \not{p}_3 \gamma_\nu (1 - \gamma^5) \not{p}_4 \right] = 256 (p_1 \cdot p_3) (p_2 \cdot p_4),$$

show that the spin-averaged matrix element is given by:

$$\frac{1}{2} \sum_{\text{spins}} |\mathfrak{M}|^2 = 64 G_F^2 (p_1 \cdot p_3) (p_2 \cdot p_4).$$

- (iv) Show that in the rest frame of the muon,

$$\frac{1}{2} \sum_{\text{spins}} |\mathfrak{M}|^2 = 32 G_F^2 m_\mu^2 |\vec{p}_2| (m_\mu - 2|\vec{p}_2|).$$

- (v) For the muon decay, it can be shown that Fermi's Golden rule leads to the expression

$$d\Gamma = \frac{\sum_{\text{spins}} |\mathfrak{M}|^2}{16(2\pi)^4 m_\mu} d|\vec{p}_2| \frac{d^3 \vec{p}_4}{|\vec{p}_4|^2}.$$

where the integral over  $|\vec{p}_2|$  runs from  $\frac{1}{2}m_\mu - |\vec{p}_4|$  up to  $\frac{1}{2}m_\mu$ , and over  $|\vec{p}_4|$  from 0 to  $\frac{1}{2}m_\mu$ .

Using this information, and integrating out first  $|\vec{p}_2|$ , and then  $d^3 \vec{p}_4 = 4\pi |\vec{p}_4|^2 d|\vec{p}_4|$ , show that the muon decay rate is given by

$$\Gamma = \frac{G_F^2 m_\mu^5}{24(2\pi)^3}.$$

- (vi) Use  $G_F = 1.166 \times 10^{-5} \text{GeV}^{-2}$  to calculate the muon life time.

**Exercise 38** *Kinematics of the Goldhaber experiment***7 Punkte**

The Goldhaber-Grodzins-Sunyar experiment was discussed during the lecture and consists of the following steps:

- $^{152}\text{Eu}$  decays by electron capture to  $^{152}\text{Sm}^*$ :  $^{152}\text{Eu} + e^- \rightarrow ^{152}\text{Sm}^* + \nu_e$ .
- $^{152}\text{Sm}^*$  deexcites by emitting a photon:  $^{152}\text{Sm}^* \rightarrow ^{152}\text{Sm} + \gamma_1$ .
- The emitted photon is absorbed by another Sm nucleus:  $\text{Sm} + \gamma_1 \rightarrow \text{Sm}^*$ .
- This  $\text{Sm}^*$  nucleus immediately deexcites to another photon  $\gamma_2$  which is captured in the NaI detector.

The key point of the experiment is that  $\gamma_1$  can only be re-absorbed by forward  $\gamma_1$  emission of by  $\text{Sm}^*$  in flight due to the recoil energy of the nuclei both in the emission and absorption processes.

- Show that  $\gamma_1$  cannot be reabsorbed by Sm if produced in an  $\text{Sm}^*$  decay at rest.
- Show that  $\gamma_1$  can be absorbed by Sm if emitted in forward direction by  $\text{Sm}^*$  in flight.

You can use the following:

- The energy released in the transition  $^{152}\text{Eu} \rightarrow ^{152}\text{Sm}^*$  is 911 keV.
- The resonance energy in question (energy difference between the two Sm levels) is 963 keV.
- The resonance width is about 1 eV (note that the natural width is only about 20 meV - the value given is after Doppler broadening due to thermal motion).
- You can approximate the neutrino and photon momentum by the energies released in the transitions in which they are produced (explain why!).
- For the energy-momentum dependence of the nuclei, you can use non-relativistic expressions (explain why!).
- Remember that you need the  $\gamma_1$  energy in the rest frame of the Sm nucleus to judge whether absorption will take place or not. You should first show that the relation between the photon energy in the frame of the source,  $E_0$ , and its energy in another frame (e.g. the lab frame),  $E$ , given by a simple Lorentz transformation, is  $\frac{E}{E_0} = \sqrt{\frac{1+\beta}{1-\beta}}$  with  $v = \beta c$  being the relative speed of the source with respect to the other frame.

**Exercise 39** *Tau lepton life time***3 Punkte**

The muon life time was calculated in exercise 37. Adjust the equation for the width given in exercise 37 (iv) to estimate the tau lepton life time, assuming it only decays leptonically. Compare the calculated value to the literature value ( $\tau_\tau = 2.91 \times 10^{-13}\text{s}$ ).

Hints:

You can neglect the muon mass in comparison to the tau lepton mass.

The calculated value disagrees with the literature value by a factor of about 2.5 which can be explained by the fact that the branching ratio of the tau lepton to electron or neutrino is actually only about 2/5.